

Abbott et al. 2016, PRL, 116, 061102

人类已经开始用引力波“聆听”宇宙了！

致密星

全号

2016年3月

前言

- 分享对致密星和高能天体物理的理解
- ppt+板书的方式
- 考试采用文献阅读报告的形式
- 2016-2020每年讲授一次

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参考书

1. **Rosswog S., Bruggen M., 2011, Introduction to high energy astrophysics**
2. Longair M. S., 2011, High energy astrophysics
3. 徐仁新, 2006, 天体物理导论
4. Shapiro S. L., Teukolsky S. A., 1983, Black hole, white dwarfs, and neutron stars
5. Frank J., et al., 2002, Accretion power in astrophysics
6. Hartle J. B., 2008, Gravity
7. Grupen C., et al., 2005, Astroparticle physics
8. Perkins D. H. 著, 来小禹等译, 2015, 粒子天体物理

1. 狭义相对论

- Rosswog & Bruggen, 2011, Introduction to high energy astrophysics, Chapter 1

1.1 引言

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1.5 狭义相对论的电动力学

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1.1 引言

- 动力学理论：
 1. 描述系统的物理量
 2. 物理量随时间的演化
- 量子力学，牛顿力学，狭义相对论

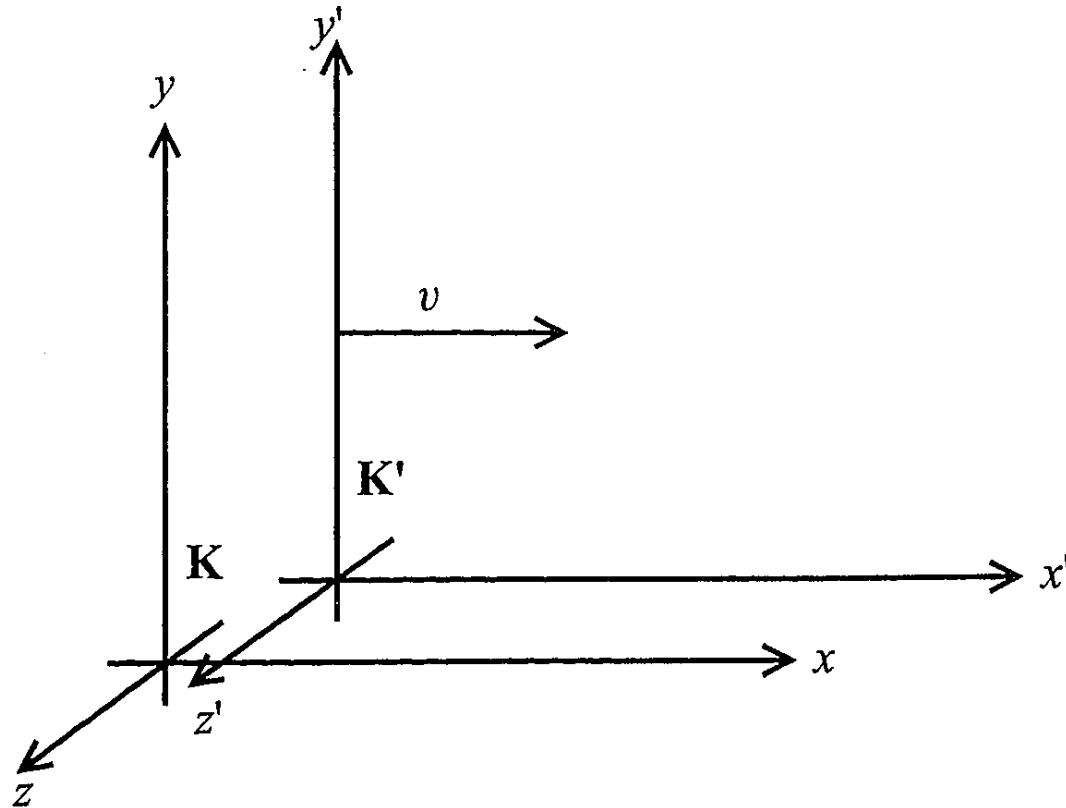


Figure 1.1 Coordinate systems used for the Lorentz transformations: the x -axes are aligned, and the relative velocity between the two frames is v .

1.2 Lorentz变换

- Lorentz 因子: $\gamma = \frac{1}{\sqrt{1 - v^2 / c^2}}$
- Lorentz 变换:
$$\begin{cases} t' = \gamma(t - \frac{vx}{c^2}) \\ x' = \gamma(x - vt) \\ y' = y \\ z' = z \end{cases}$$

1.3 狭义相对论效应

- Lorentz变换是关于时空事件的变换，时间，长度，速度等物理量可由Lorentz变换推导得出

1.3.1 长度收缩

1.3.2 时间膨胀

1.3.3 相对论集束效应(beaming)

1.3.4 Doppler 效应和视超光速(superluminal motion)

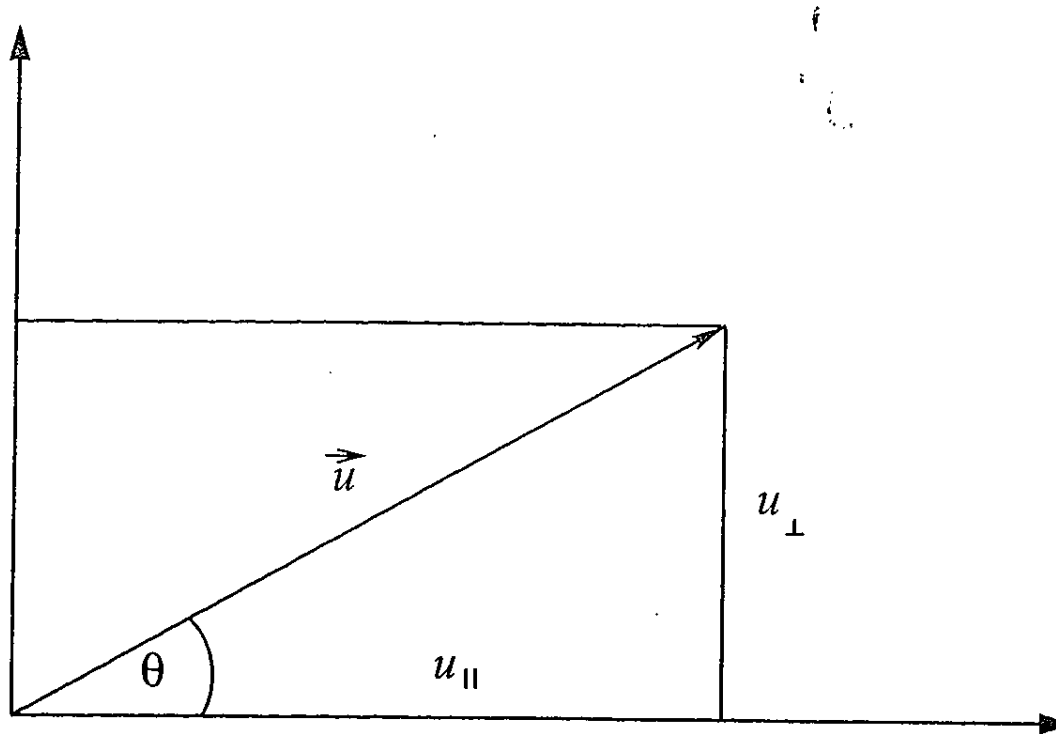


Figure 1.2 For the special relativistic transformation of velocities: splitting the velocity \vec{u} into a component parallel (u_{\parallel}) and perpendicular (u_{\perp}) to the velocity between the two frames, v .

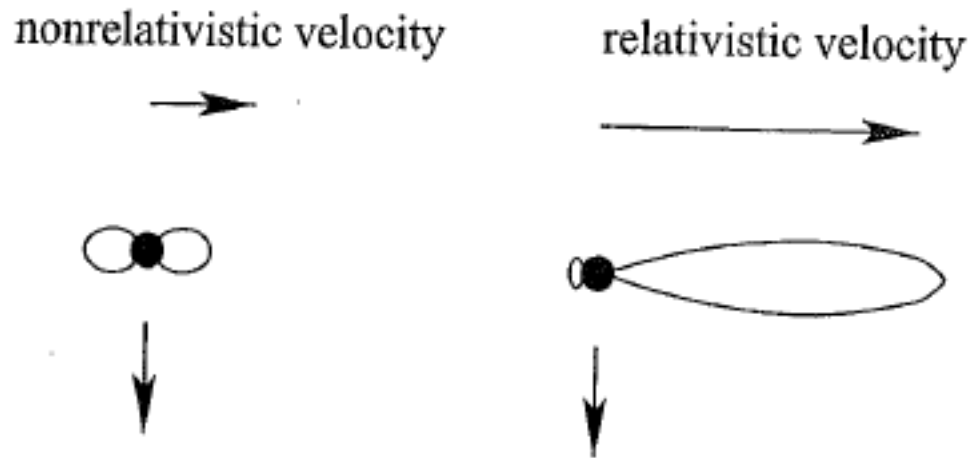


Figure 1.4 Change of a dipole pattern (acceleration perpendicular to the indicated velocity) due to relativistic beaming (left: nonrelativistic, right: relativistic velocity).

Rosswog & Bruggen 2011

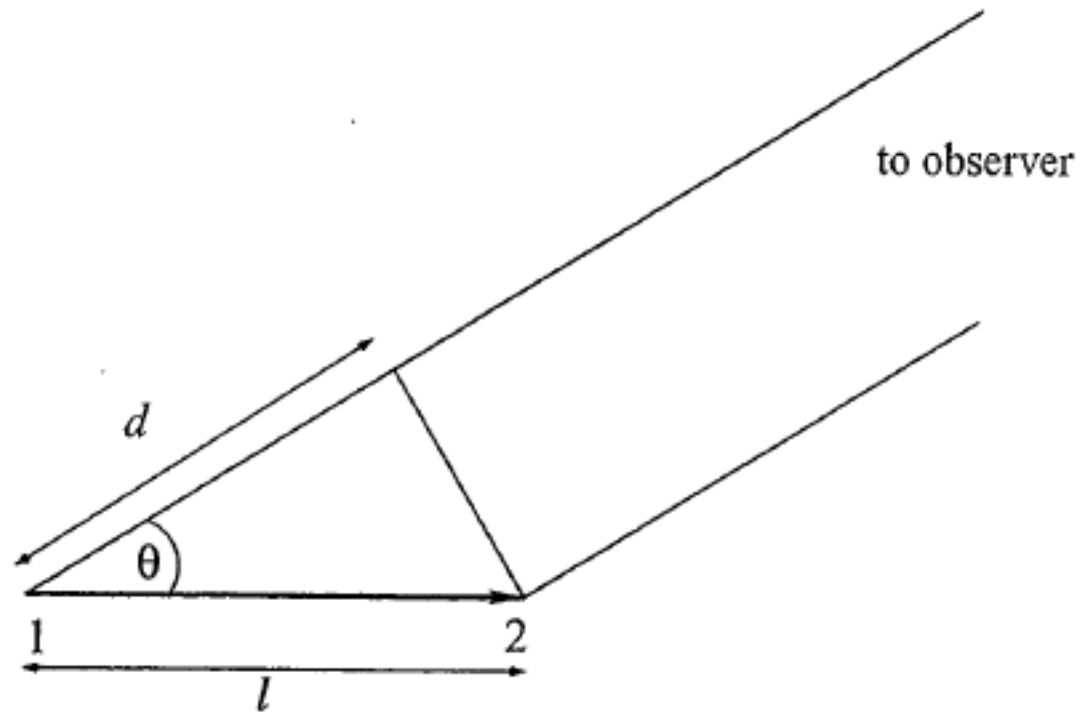


Figure 1.5 A rapidly moving, periodically emitting source travels during one period from point 1 to 2.

1.4 四矢量

- 三维空间中的转动
- 四维时空中的转动

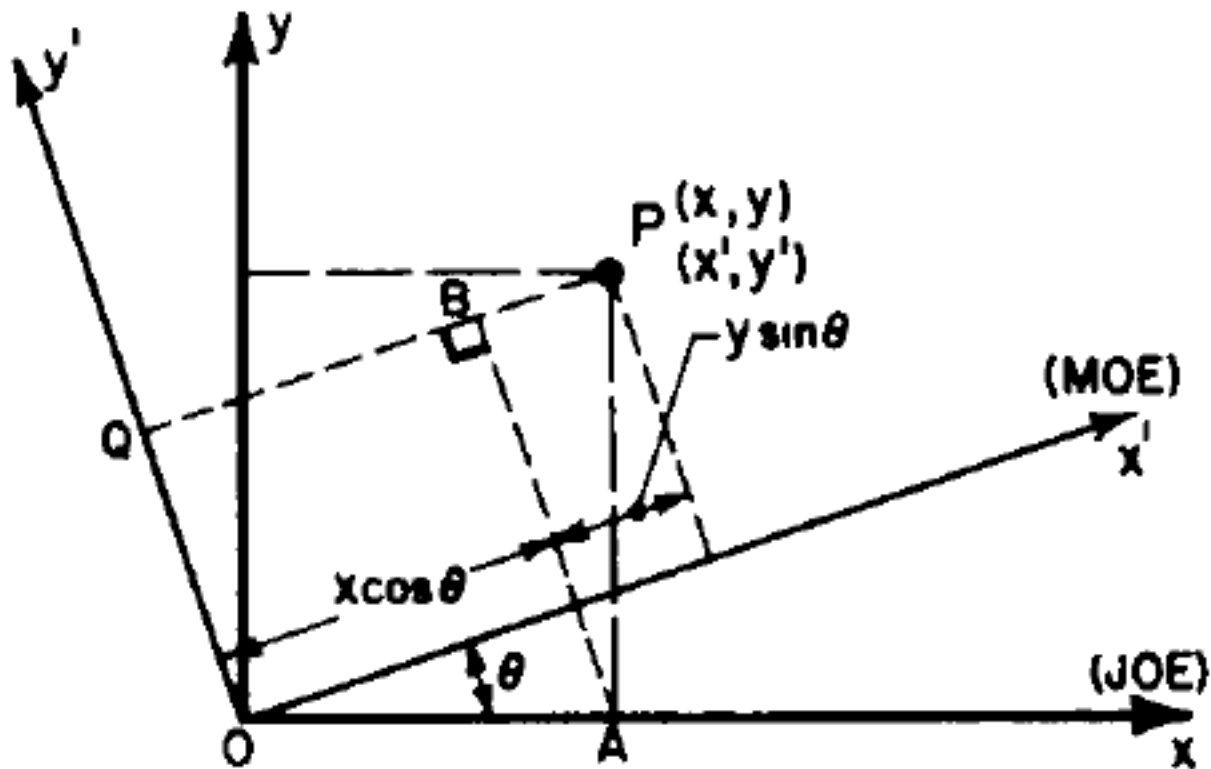


Fig. 11-2. Two coordinate systems having different angular orientations.

1.5 狭义相对论的电动力学

- 电动力学自然满足Lorentz变换

1.6所有的经典物理

(Feynman物理学讲义第2卷 第18章Maxwell方程组)

Table 18-1 Classical Physics

Maxwell's equations	
I. $\nabla \cdot \mathbf{E} = \frac{\rho}{\epsilon_0}$	(Flux of \mathbf{E} through a closed surface) = (Charge inside)/ ϵ_0
II. $\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$	(Line integral of \mathbf{E} around a loop) = $-\frac{d}{dt}$ (Flux of \mathbf{B} through the loop)
III. $\nabla \cdot \mathbf{B} = 0$	(Flux of \mathbf{B} through a closed surface) = 0
IV. $c^2 \nabla \times \mathbf{B} = \frac{\mathbf{j}}{\epsilon_0} + \frac{\partial \mathbf{E}}{\partial t}$	c^2 (Integral of \mathbf{B} around a loop) = (Current through the loop)/ ϵ_0 + $\frac{\partial}{\partial t}$ (Flux of \mathbf{E} through the loop)
<div style="border: 1px solid black; padding: 5px; display: inline-block;"> <p>Conservation of charge</p> <p>$\nabla \cdot \mathbf{j} = -\frac{\partial \rho}{\partial t}$ (Flux of current through a closed surface) = $-\frac{\partial}{\partial t}$ (Charge inside)</p> </div>	
Force law	
$\mathbf{F} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B})$	
Law of motion	
$\frac{d}{dt}(\mathbf{p}) = \mathbf{F}$, where $\mathbf{p} = \frac{m\mathbf{v}}{\sqrt{1 - v^2/c^2}}$ (Newton's law, with Einstein's modification)	
Gravitation	
$\mathbf{F} = -G \frac{m_1 m_2}{r^2} \mathbf{e}_r$	

思考题

1. 如何有光速不变导出Lorentz变换(力学)?
2. 用学到的知识考虑Compton 散射(Rosswog & Bruggen 1.4.6 / 原子物理等)。